# The Infrared Spectrum of Acetylene- $d_{1}{ }^{1}$ 

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The vibrational-rotational spectrum of acetylene- $d_{1}$ has been observed and measured from 1800 to $9200 \mathrm{~cm} .^{-1}$ Rotational structure sufficient to enable an analysis was found for 28 bands. From the average combination-difference values of the 14 best bands the rotational constants were found to be, $B_{0}=0.99169 \pm 0.00015 \mathrm{~cm} .^{-1}$ and $D_{0}=1.08 \pm 0.26 \times$ $10^{-6} \mathrm{~cm} .^{-1}$ The vibrational energy system is fully discussed. It is pointed out that the vibrational energy levels cannot be fit by the usual quadratic expansion of the energy in terms of the vibrational quantum numbers. Evidence is presented to show that at least cubic terms must be included in the expansion.

## Introduction

Although the infrared spectra of acetylene and acetylene- $d_{2}$ have been extensively investigated under high resolution, ${ }^{2-6}$ no high resolution work has been reported on the partially deuterated species in the non-photographic infrared. The vibrational energy systems of $\mathrm{C}_{2} \mathrm{H}_{2}$ and $\mathrm{C}_{2} \mathrm{D}_{2}$ are still not well understood in spite of the extensive work. Since $\mathrm{C}_{2} \mathrm{HD}$ is less symmetric, there are no vibrational transitions which are forbidden by symmetry and hence there is a better chance to observe enough bands to enable an understanding of its energy system and perhaps furnish some clues which will help unravel the other two molecules. The spectrum of $\mathrm{C}_{2} \mathrm{HD}$ has been observed from 1800 to $9200 \mathrm{~cm} .^{-1}$. In this region 28 bands were observed for which rotational analyses could be made. In addition several Q branches were measured for which the $P$ and R branch absorption was too weak to be observed.

## Experimental

In the spectral region from 1 to $5.5 \mu$, there have been observed 28 bands of $\mathrm{C}_{2} \mathrm{HD}$ for which the centers have been determined. These bands varied considerably in intensity


Fig. 1.-The $\nu_{1}+\nu_{3}$ and the $\nu_{1}+\nu_{3}+\nu_{4}^{1}-\nu_{4}^{1}$ bands of $\mathrm{C}_{2} \mathrm{HD}$ in the region of $5900 \mathrm{~cm} .^{-1}$. The rotational quantum numbers are given below the trace for the $\nu_{1}+\nu_{2}+\nu_{4}^{1}$ band. A ten meter cell with a pressure of 16 cm . ( Hg ) of $\mathrm{C}_{2} \mathrm{HD}$ was used for this region.

[^0] 45, 102 (1955).
and the measurements were made on the gas in two cells of 1 m . and 10 m . in length and at pressures from a few millimeters to 76 cm . ( Hg ).
Two instruments were used in obtaining the data. One instrument employs a grating of 15,000 lines $/ \mathrm{in}$. and a PbS detector. Other details of this spectrometer have been previously published. ${ }^{7}$ The second instrument contained a 7500 lines/in. grating and a cooled PbTe cell and it was used for the observations in the region from 3 to $5.5 \mu$. A detailed description of this instrument has been reported. 8
The rotational lines were measured by comparison with the fringe system from a Fabry-Perot interferometer. The absorption spectrum and the fringe system were recorded simultaneously on a two-pen recorder. The fringe system was calibrated by superimposing well known spectral lines on the chart at intervals throughout the band. This method of measurement is well suited for high precision measurements. When bands are not seriously overlapped by other bands (hot bands and isotopic bands) the band centers can be determined to a few hundredths of a wave number.

The $\mathrm{C}_{2} \mathrm{HD}$ gas was made by Dr. Hellman of this Bureau. The gas was analyzed by a mass spectrometer and showed $16.8 \% \mathrm{C}_{2} \mathrm{H}_{2}, 47.8 \% \mathrm{C}_{2} \mathrm{HD}$ and $35.4 \% \mathrm{C}_{2} \mathrm{D}_{2}$. The spectra of the three components caused serious overlapping only in the region of $6600 \mathrm{~cm} .^{-1}$.
The results obtained in two regions of the spectrum are shown in Figs. 1 and 2. White recorder paper was used so that photographs could be made of the actual tracing from the instrument. Figure 1 shows the bands $\nu_{1}+\nu_{3}$ and $\nu_{1}+$ $\nu_{3}+\nu_{4}^{1}-\nu_{4}^{1}$ in the region of $5900 \mathrm{~cm} .^{-1}$. 2580 and shows the three bands $\nu_{3}, \nu_{3}+\nu_{4}^{2}-\nu_{4}^{1}$, and $\nu_{3}+\nu_{5}^{\frac{1}{5}}-\nu_{5}^{2}$.

Rotational Analysis.-The bands were analyzed by first assigning. $J$ values to the observed absorption peaks. An equation for each $J$ value in a given band was then obtained of the form
$\nu=\nu_{0}+\left(B^{\prime}+B^{\prime \prime}\right) m+\left(B^{\prime}-B^{\prime \prime}\right) m^{2}+$ $4 \mathrm{Dm}^{3}$ (1)
where $m=J+1$ for the R-branch and $m=-J$ for the P-branch. The set of equations for each band was solved by the method of least squares to obtain values for $\nu_{0}, B^{\prime}, B^{\prime \prime}$ and $D$. The least squares reductions were carried out on SEAC, the NBS electronic computer. The observed absorption frequencies and rotational assignments for several of the stronger bands are given in Table I.

The value of $B_{0}$ varied somewhat from band to band. This is believed to be due to the influence of overlapping hot bands. In order to get a best value
(7) N. Gailar and E. K. Plyler ibid., 46, 392 (1952).
(8) E. K. Plyler, L. R. Blaine and W. S. Connor, J. Opt. Soc. Am.,

Table I
Observed Frequencies and Rotational Assignments for Several $\mathrm{C}_{2} \mathrm{H}$ D Bands

| $T^{\nu 2}=1853.75$ |  | $R^{\boldsymbol{\nu} / 2}=2583.638$ |  | $R^{\nu} \nu_{1}=3335.58$ |  | ${ }_{R}{ }_{1}+\nu_{6}{ }^{1}=3995.77{ }_{P}$ |  | ${ }_{R}^{\nu_{2}}+v_{2}=4415.882$ |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1855.775 |  | 2585.575 |  | 3337.40 |  | 3997.72 |  | 4417.86 |  |
| 1857.737 | 1851.807 | 2587.542 | 2581.747 | 3339.37 | 3333.54 | 3999.69 |  | 4419.75 | 4413.89 |
| 1859.674 | 1849.807 | 2589.492 | 2579.688 | 3341.39 | 3331.43 | 4001.68 |  | 4421.68 | 4411.89 |
| 1861.597 |  | 2591.426 | 2577.678 | 3343.24 | 3329.64 | 4003.61 | 3989.81 | 4423.60 | 4409.85 |
| 1863.570 |  | 2593.344 | 2575.639 | 3345.27 | 3327.64 | 4005.55 | 3987.78 | 4425.51 | 4407.80 |
| 1865.500 |  | 2595.234 | 2573.581 | 3347.32 | 3325.68 | 4007.49 | 3985.75 | 4427.35 | 4405.70 |
| 1867.486 | 1841.773 | 2597.128 | 2571.542+ | 3349.05 | 3323.53 | 4009.35 | 3983.72 | 4429.15 | 4403.64 |
|  | 1839.731 | 2598.996 | 2569.517+ | 3351.11 | 3321.55 | 4011.26 | 3981.65 | 4430.99 | 4401.52 |
| 1871.258 | 1837.687 | 2600.898 | 2567.380 | 3352.99 | 3319.50 | 4013.15 | 3979.63 | 4432.75 | 4399.39 |
| 1873.163 | 1835.646 | 2602.735 | 2565.297 | 3354.87 | 3317.48 | 4015.03 | 3977.52 | 4434.51 | 4397.26 |
| 1875.051 | 1833.573 | 2604.561 | 2563.214 | 3356.83 | 3315.38 | 4016.91 | 3975.48 | 4436.27 | 4395.05 |
| 1876.924 | 1831.491 | 2606.379 | 2561.048 | 3358.55 | 3313.31 | 4018,76 | 3973.39 | 4437.96 | 4392.87 |
| 1878.789 |  | 2608.222 | 2558.976 | 3360.31 | 3311.20 | 4020.59 | 3971.32 | 4439.70 | 4390.65 |
| 1880.654 | 1826.351 | 2610.037 | 2556.816 | 3362.30 | 3309.08 | 4022.45 | 3969.22 | 4441.35 | 4388.43 |
| 1882.515 |  | 2611.814 | 2554.682 | 3364.23 | 3307.08 | 4024.28 | 3967.11 | 4443.00 | 4386.12 |
| 1884.352 | 1823.143 | 2613.527 | 2552.497 | 3365.99 | 3304.92 | 4026.11 | 3965.18 | 4444.65 | 4383.81 |
|  | 1821.047 | 2615.298 | 2550.286 | 3367.89 | 3302.77 | 4027.96 |  | 4446.24 | 4381.57 |
|  |  | 2617.000 | 2548.092 | 3369.65 | 3300.55 | 4029.75 | 3960.67 | 4447.83 | 4379.21 |
|  |  | 2618.720 |  | 3371.43 | 3298.46 | 4031.56 | 3958.52 | 4449.38 | 4376.85 |
| 1891.607 | 1814.654 | 2620.452 | 2543.633 | 3373.26 | 3296.37 |  | 3956.40 | 4450.91 | 4374.48 |
| 1893.413 | 1812.497 | 2622.141 | $2541.426+$ | 3374.97 | 3294.17 | 4035.06 | 3954.20 | 4452.45 | 4372.11 |
|  |  | 2623.831 | 2539.242 | 3376.78 | 3292.01 | 4036.80 | 3952.00 | 4453.93 | 4369.74 |
| 1896.974 | 1808.22 | 2625.496 | 2536.933 | 3378.52 | 3289.92 | 4038.54 | 3949.86 | 4455.41 | 4367.27 |
| 1898.739 | 1806.046 | 2627.157 | 2534.633 | 3380.26 | 3287.71 | 4040.24 |  | 4456.88 | 4364.82 |
| 1900.500 | 1803.869 | 2628.786 | 2532.436 | 3381.94 | 3285.42 | 4041.99 | 3945.40 | 4458.36 | 4362.34 |
| 1902.252 | 1801.679 | 2630.420 | 2530.106 | 3383.68 | 3283.08 | 4043.70 |  | 4459.72 | 4359.82 |
| 1903.996 |  | 2632.035 | 2527.797 | 3385.45 | 3280.75 | 4045.40 |  | 4461.10 | 4357.32 |
| 1905.745 |  | 2633.652 | 2525.480 | 3387.14 |  | 4047.06 |  | 4462.49 | 4354.76 |
| 1907.517 |  | 2635.181 | 2523.186 | 3388.78 |  | 4048.70 |  | 4463.79 | 4352, 14 |
| 1909.182 |  | 2636.707 | 2520.851 |  |  |  |  | 4465.16 | 4349.61 |
| 1910.872 |  | 2638.360 | 2518.462 |  |  |  |  | 4466.46 |  |
| 1912.589 |  | 2639.916 | 2516.048 | 3393.80 |  |  |  | 4467.70 | 4344.35 |
| 1914.253 |  | 2641.405 | 2513.679 | 3395.47 |  |  |  | 4468.93 |  |
| 1915.248 |  |  | 2511.296 |  |  |  |  | 4470.08 |  |
|  |  |  | 2508.828 | 3398.74 |  |  |  | 4471.37 |  |
|  |  |  | 2506.430 | 3400.36 |  |  |  |  |  |
|  |  |  | 2503.996 | 3402.00 |  |  |  |  |  |
|  |  |  | 2501.556 |  |  |  |  |  |  |


| $\underset{R}{p_{1}+v_{s}{ }^{0}=4643.244}$ |  |
| :---: | :---: |
|  | 4641.22 |
| 4649.12 | 4639.23 |
| 4651.13 | 4637.30 |
| 4653.08 | 4635.32 |
| 4655.02 | 4633.23 |
| 4657.06 | 4631.24 |
| 4658.96 | 4629.26 |
| 4660.92 | 4627.27 |
| 4662.86 | 4625.27 |
| 4664.79 | 4623.23 |
| 4660.66+ | 4621.23 |
| 4668.68 | 4619.19 |
| 4670.60 | 4617.15 |
| 4672.52 | 4615.16 |
| 4674.46 | 4613.11 |
| 4676.34 | 4611.09 |
| 4678.26 | 4609.01 |
| 4680.14 | 4606.99 |
| 4682.04 | 4604.93 |
| 4683.91 | 4602.93 |
| 4685.75 | 4600.81 |
| 4687.68 | 4598.74 |
| 4689.56 | 4596.65 |
|  | 4594.53 |

Table I (Continued)

| $\nu_{2}+\nu_{J}+\nu_{4}=$ | 4915.719 |
| :--- | :--- |
|  |  |
|  |  |
|  |  |
| 4923.63 |  |
| 4925.41 | 4907.73 |
| 4927.22 | 4905.75 |
| 4929.06 | 4903.55 |
| 4930.97 | 4901.47 |
| 4932.83 | 4897.33 |
| 4934.49 | 4895.02 |
| 4936.32 | 4892.89 |
| 4938.00 | 4890.62 |
| 4939.66 | 4888.31 |
| 4941.34 | 4886.18 |
| 4943.01 | 483.89 |
| 494.63 | 481.62 |
| 4946.37 | 4879.40 |
| 4948.00 | 4877.01 |
| 4952.56 | 4874.48 |
| 4954.14 | 4871.79 |
|  |  |
|  |  |

$R^{2 \nu_{2}=5123.114}{ }_{P}$
5125.08
5126.96 $5128.92 \quad 5119.12$ 5130.74
5134.48
5136.22
5138.00
5139.75
$\begin{array}{ll}5141.48 & 5106.56 \\ 5143.20 & 5104.32\end{array}$
$5143.20 \quad 5102.13$
$\begin{array}{ll}5144.84 & 5099.76 \\ 5146.46 & 5097.63\end{array}$
$\begin{array}{ll}5146.46 & 5097.63 \\ 5148.08 & 5095.22\end{array}$
$\begin{array}{ll}5148.08 & 5095.22 \\ 5149.65 & 5092.89\end{array}$
$\begin{array}{ll}5149.65 & 5092.89 \\ 5151.18 & 5090.44\end{array}$
$5152.68 \quad 5088.22$

| 5154.19 | 5085.83 |
| :--- | :--- |
|  | 5083.27 |

5157.04
5158.49
$5159.98 \quad 5078.40$
$\begin{array}{ll}5159.98 & 5075.92 \\ 5161.30 & 5073.38\end{array}$
$\begin{array}{ll}5161.30 & 5073.38 \\ 5162.64 & \end{array}$
$\begin{array}{ll}5162.64 & 5068.16\end{array}$
$5165.18+\quad 5065.64$
5166.42
$5167.64+\quad 5060.16$
$5168.89+\quad 5057.49$
$5170.09 \quad 5054.91$
$5171.12 \quad 5052.16$
5172.28
5174.40
5175.46
$\nu_{1} \underset{R}{\nu_{2}}+\nu_{3}{ }^{1}=\underset{P}{p}$
5798.93
5800.83
$\begin{array}{ll}5802.69+ & \\ 5804.51 & 5786.87\end{array}$
$5806.31 \quad 5784.73$
$5808.12 \quad 5782.73$
$5809.87 \quad 5780.58$
$5811.62 \quad 5778.42$
$5813.34 \quad 5776.20$
$5815.02 \quad 5773.94$
$5816.63 \quad 5771.73$
$5818.28 \quad 5769.41$
$\begin{array}{ll}5819.88 & 5767.14 \\ 5821.47 & 5764.79\end{array}$
$\begin{array}{ll}5823.00 & 5764.79 \\ 5762.40\end{array}$
$5824.50 \quad 5759.98$
$5825.94 \quad 5757.59$
$\begin{array}{ll}5827.46 & 5755.19 \\ 5828.81 & 5752.70\end{array}$
$5828.81 \quad 5752.70$



Fig. 2.-The $\nu_{3}, \nu_{3}+\nu_{4}^{1}-\nu_{4}^{1}$ and $\nu_{3}+\nu_{5}^{1}-\nu_{5}^{1}$ bands of $\mathrm{C}_{2} \mathrm{HD}$ in the region of $2580 \mathrm{~cm} .^{-1}$. The rotational quantum numbers for $\nu_{3}$ are marked below the spectrum. A one meter cell with a pressure of 1.5 cm . ( Hg ) was used for observing this band.

Table II

| $\nu_{0}+B^{\prime} l^{\prime 2}+B^{\prime \prime} l^{\prime \prime 2}$ | $\nu_{1}$ | 2 |  | Infrared Spectrum of $\mathrm{C}_{2} \mathrm{HD}$ |  |  |  |  |  |  | $\left(B_{\text {obsd. }}^{\prime}-B^{\prime \prime}\right)$ | $\left(\begin{array}{l} \left(B^{\prime}-B^{\prime \prime}\right) \\ \text { calcd. } \end{array}\right.$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  |  |  | $9_{s} 1$ | $v_{s}{ }^{1}$ | $\nu_{1}$ | ${ }_{72}$ | $\underset{\eta_{i}}{\substack{\text { Lower }}}$ | $v_{4}{ }^{1}$ | 0 |  |  |
| 1200.63 |  |  |  | $1{ }^{1}$ | $1{ }^{1}$ |  |  |  |  |  | 0.0050 |  |
| 1850.177 |  |  |  |  |  |  |  |  |  |  | . 00432 |  |
| 1852.46 |  | 1 |  |  | $1{ }^{1}$ |  |  |  |  | $1^{1}$ | . 0038 | 0.0042 |
| 1853.75 |  | 1 |  |  |  |  |  |  |  |  | . 00422 | 0.00422 |
| 2065.24 |  |  | 1 |  |  |  |  |  | $1{ }^{1}$ |  | . 00699 |  |
| 2568.29 |  |  | 1 | $1^{1}$ |  |  |  |  | $1{ }^{1}$ |  | . 00645 | 0.00670 |
| 2580.81 |  |  | 1 |  | $1{ }^{1}$ |  |  |  |  | $1{ }^{1}$ | . 00677 | 0.00670 |
| 2583.638 |  |  | 1 |  |  |  |  |  |  |  | . 00670 | 0.00670 |
| 3086.61 |  |  | 1 | $1^{1}$ |  |  |  |  |  |  | . 00614 |  |
| 3335.58 | 1 |  |  |  |  |  |  |  |  |  | . 00480 | 0.00480 |
| 3995.777 | 1 |  |  |  | $1{ }^{1}$ |  |  |  |  |  | . 00508 |  |
| 4362.11 | 1 |  |  | $2^{0}$ |  |  |  |  |  |  | . 00128 |  |
| 4385.76 |  | 2 |  |  | $1{ }^{1}$ |  |  |  |  |  | . 0003 |  |
| 4397.47 |  | 1 | 1 | $1{ }^{1}$ |  |  |  |  | $1{ }^{1}$ |  | . 0108 | 0.0109 |
| 4415.882 |  | 1 | 1 |  |  |  |  |  |  |  | . 01091 | 0.01092 |
| 4513.57 | 1 |  |  | $1{ }^{1}$ | $1{ }^{1}$ |  |  |  |  |  | . 00092 |  |
| 4643.244 | 1 |  |  |  | $2^{0}$ |  |  |  |  |  | . 00198 |  |
| 4915.719 |  | 1 | 1 | $1^{1}$ |  |  |  |  |  |  | . 00941 |  |
| 5123.114 |  |  | 2 |  |  |  |  |  |  |  | . 01338 | 0.01340 |
| 5610.03 |  |  | 2 | $1^{1}$ |  |  |  |  |  |  | . 01390 |  |
| 5795.014 | 1 | 1 |  |  | $1{ }^{1}$ |  |  |  |  |  | . 01372 |  |
| 5894.15 | 1 |  | 1 | $1{ }^{1}$ |  |  |  |  | $1{ }^{1}$ |  | . 0113 | 0.0115 |
| 5913.312 | 1 |  | 1 |  |  |  |  |  |  |  | . 01143 | . 01150 |
| 6564.24 | 2 |  |  |  | $1^{1}$ |  |  |  |  | $1^{1}$ | . 0119 | . 0960 |
| 6569.36 | 2 |  |  |  |  |  |  |  |  |  | . 01221 | . 0960 |
| 6932.20 |  | 1 | 2 |  |  |  |  |  |  |  | . 01755 | . 01762 |
| 7620.34 |  |  | 3 |  |  |  |  |  |  |  | . 02016 | . 02010 |
| 8409.18 | 2 | 1 |  |  |  |  |  |  |  |  | . 01386 | . 01382 |
| 9139.02 | 2 |  | 1 |  |  |  |  |  |  |  | . 01617 | . 01630 |
| $\begin{gathered} B_{0}=0.99169 \pm 0.00015 \mathrm{~cm} .^{-1} \\ D_{0}=1.08 \pm 0.26 \times 10^{-6} \mathrm{~cm} .^{-1} \end{gathered}$ |  |  |  |  |  |  |  |  |  |  |  |  |

for $B_{0}$, use was made of the combination difference relation

$$
\begin{align*}
& \Delta F_{2}{ }^{\prime \prime}= R(J-1)- \\
& P(J+1)=  \tag{2}\\
& 4 B^{\prime \prime}(J+1 / 2)+8 D^{\prime \prime}(J+1 / 2)^{3}
\end{align*}
$$

The $\Delta F_{2}$ values were obtained by averaging the $\Delta F_{2}$ values of the 14 best bands. These equations were then solved by the method of least squares. The values of the rotational constants obtained in this manner are, $B_{0}=0.99169 \pm 0.00015 \mathrm{~cm} .^{-1}$ and $D_{0}=1.08 \pm 0.26 \times 10^{-6} \mathrm{~cm} .^{-1}$. The values of $\nu_{0}$
and $B^{\prime}-B^{\prime \prime}$ obtained from the solutions of eq. 1 are given in Table II. From the analyses of $\nu_{1}, \nu_{2}$ and $\nu_{3}$ one obtains directly values of $\alpha_{1}, \alpha_{2}$ and $\alpha_{3}$ in the expansion of the inertial constants in terms of the vibrational constants

$$
\begin{equation*}
B_{v}=B_{e}+\sum_{i=1}^{5} \alpha_{i}\left(V_{i}+d_{i} / 2\right) \tag{3}
\end{equation*}
$$

The values of the constants are $\alpha_{1}=0.00480 \mathrm{~cm} .^{-1}$, $\alpha_{2}=0.00422$ and $\alpha_{3}=0.00670$. These values can be checked by calculating $B_{v}$ for the states in which
only $v_{1}, v_{2}$ and $v_{3}$ are greater than zero. This has been done and the results are included in the last column of Table II. It is readily seen that the agreement is excellent in all cases except $2 \nu_{1}$ at $6569.36 \mathrm{~cm}^{-1}$. The reason for this is not immediately apparent. As we shall see later this band is also a misfit in the vibrational analysis. The $B_{\text {e }}$ value cannot be determined until adequate values of $\alpha_{4}$ and $\alpha_{5}$ are obtained. These values cannot be obtained from the present work since no Q-branches were resolved and no l-type doubling was observed. However $B_{0}$ values have been determined in this Laboratory for $\mathrm{C}_{2} \mathrm{H}_{2}$ and $\mathrm{C}_{2} \mathrm{D}_{2}$ as well as $\mathrm{C}_{2} \mathrm{HD}$. The values were determined from the average $\Delta F_{2}$ values of many bands and a least squares reduction of the data. The results are tabulated in Table III along with the bond distances which one can calculate from these inertial constants.

Table III

| Molecular Constants of Acetylene ${ }^{a}$ |  |
| :--- | :---: | ---: |
|  | $B_{0}$, cm..$^{-1}$ |$]$.

a The atomic constants used are taken from the recent work of J. A. Bearden and John S. Thomsen "A Survey of Atomic Constants," a report prepared for Bureau of Ordnance, U. S. Navy, 1955.

Vibrational Analysis.-The vibrational energy levels of a molecule are usually expanded in terms of the vibrational quantum numbers in an equation of the form

$$
\begin{align*}
& E_{\mathrm{v}}-E_{0}= \sum_{i} \omega_{i} V_{i}-\sum_{i_{i} k \geq i} X_{i k} V_{i} V_{k}- \\
& \sum_{i j k} Y_{i j k} V_{i} V_{j} V_{k}+\ldots \tag{4}
\end{align*}
$$

In most spectra studied so far, it has been sufficient to neglect terms above those quadratic in the vibrational quantum numbers. For $\mathrm{C}_{2} \mathrm{HD}$ this approximation seems to be inadequate. Consider $\nu_{1}$ and its overtones as is done in Table IV. The first, second and third differences are tabulated. If this series could be represented adequately by a quadratic expression, the second differences would be equal. If this series could be adequately represented by a cubic expression, the third differences would be equal. Clearly neither of these conditions is fulfilled. However if $2 \nu_{2}$ were about $1 \mathrm{~cm} .^{-1}$ higher, then the third differences could be equal and a cubic expression would suffice. The $\mathrm{C}_{2} \mathrm{HD}$ band at 6569 is badly overlapped by both $\mathrm{C}_{2} \mathrm{H}_{2}$ and $\mathrm{C}_{2} \mathrm{D}_{2}$ as well as hot bands of $\mathrm{C}_{2} \mathrm{HD}$. It is conceivable that the analysis here is in error, which would help straighten things out. One can also consider the series $n \nu_{1}+\nu_{3}$ as is done in the second group of Table IV. Again we clearly see that a quadratic term is not adequate. However if one uses the $\Delta_{3} \nu$ to calculate $y_{111}$, a set of constants $\omega_{1}^{\circ}, X_{11}$ and $Y_{111}$ can be obtained which faithfully calculates the
series $n \nu_{1}$ except for $n=2$ which it predicts about $1 \mathrm{~cm} .^{-1}$ higher than it is observed. Further, if cubic cross terms were not important, then the $\Delta_{2} \nu$ values for the two series should be equal, a condition which, again, is not fulfilled.

Table IV

| Table Showing Band Series Regularities |  |  |  |  |
| :---: | :---: | :---: | :---: | :---: |
| $v_{1} v_{2} v_{s} v_{0} v_{0}{ }^{1}$ | $\nu, \mathrm{cm} .^{-1}$ | $\Delta \nu$ | $\Delta_{2} \nu$ | $\Delta 8 \nu$ |
| 00000 | 0 |  |  |  |
|  |  | 3335.58 |  |  |
| 10000 | 3335.58 |  | 101.80 |  |
|  |  | 3233.78 |  | 5.06 |
| 20000 | 6569.36 |  | 96.74 |  |
|  |  | 3137.04 |  | 0.10 |
| 30000 | 9706.4 |  | 96.64 |  |
|  |  | 3040.40 |  |  |
| 40000 | 12746.8 |  |  |  |
| 00100 | 2583.64 |  |  |  |
|  |  | 3329.67 |  |  |
| 10100 | 5913.31 |  | 103.96 |  |
|  |  | 3225.71 |  | 2.23 |
| 20100 | 9139.02 |  | 101.73 |  |
|  |  | 3123.98 |  |  |
| 30100 | 12263.0 |  |  |  |
| 00000 | 0 |  |  |  |
|  |  | 2583.64 |  |  |
| 00100 | 2583.64 |  | 44.21 |  |
|  |  | 2539.43 |  | 2.07 |
| 00200 | 5123.07 |  | 42.14 |  |
|  |  | 2497.29 |  |  |
| 00300 | 7620.34 |  |  |  |
| 01000 | 1853.75 |  |  |  |
|  |  | 2562.13 |  |  |
| 01100 | 4415.88 |  | 45.81 |  |
|  |  | 2516.32 |  | 2.08 |
| 01200 | 6932.20 |  | 43.72 |  |
|  |  | 2472.60 |  |  |
| 01300 | 9404.8 |  |  |  |

As further evidence the data and differences for the two series $n \nu_{3}$ and $\nu_{2}+n \nu_{3}$ are also given in Table II. These series cannot be followed to sufficiently high values of $v_{3}$ to determine whether or not the third differences are constant. Certainly the third differences are the same for each series, hence an unambiguous value of $Y_{333}$ can be determined. Again the respective values of $\Delta_{2} \nu$ in the two series are not equal indicating a sizable contribution from such terms as $Y_{122}$ and $Y_{122}$. This evidence points quite conclusively to the fact that at least terms cubic in the vibrational quantum numbers are necessary to represent the vibrational energy levels. Several constants have been determined and are given in Table $V$.

Table V
Vibrational Constants of $\mathrm{C}_{2} \mathrm{HD}$

$$
\begin{array}{crl}
\omega_{1}^{\circ}=3386.92 \mathrm{~cm} . & X_{11}=51.77 & Y_{111}=-0.43 \\
\omega_{3}^{\circ}=2606.45 & X_{33}=23.16 & Y_{333}=-0.35 \\
X_{23}+Y_{223}=20.71 & Y_{233}=0.80 \\
X_{13}+Y_{133}=4.35 & Y_{113}=1.56
\end{array}
$$

Constants from levels in which all $\nu_{i}<2$

$$
\begin{gathered}
X_{23}=21.51 \mathrm{~cm}^{-1} \quad X_{24}=3.16 \quad X_{34}=15.26 \\
\nu^{1}{ }_{4}=518.26 \mathrm{~cm}^{-1}
\end{gathered}
$$

One might suggest a resonance phenomenon as the reason for this unexpected situation. It is felt that this is ruled out by the regularities observed in the second and third differences in Table III. Even a small resonance interaction would introduce serious irregularities. It is possible that $2 \nu_{1}$ may be displaced due to some local resonance interaction, but the other resonating level is not obvious. If the transition involving the other level is normally too weak to be seen, one would expect it to borrow intensity enough from $2 \nu_{1}$ to be seen, for $2 \nu_{1}$ is a rather strong band.

Using levels in which no normal mode is excited by more than one quantum of energy, it is possible to evaluate, in several ways, what appear to be constants in a quadratic energy expression. In the light of the information in Table III this is fortuitous. Since no level was excited by $v>1$, the resulting constants are probably combinations of $X_{i j}$ 's and $Y_{i j k}$ 's. Constants which can be evaluated in this way are given in Table IV where the designa-
tion used is that appropriate to a quadratic equation.

A final piece of information which can be deduced is the band center of $\nu_{4}^{1}$. This can be determined three different ways: $\nu_{3}$ and $\nu_{3}-\nu_{4}^{0}, \nu_{2}+\nu_{3}+\nu_{4}^{1}$ and $\nu_{2}+\nu_{3}+\nu_{4}^{1}-\nu_{4}^{1}$, and $\nu_{3}+\nu_{4}^{1}$ and $\nu_{3}+\nu_{4}^{1}-\nu_{4}$. The values obtained for $\nu_{4}^{1}$ are $518.40,518.25$ and $518.32 \mathrm{~cm} .^{-1}$. Since there is no reason to prefer any one of these values, the average $518.31 \mathrm{~cm}^{-1}$ should represent the band center satisfactorily. The poor agreement can be explained by the fact that these values were deduced from hot bands which were badly overlapped by a strong $\mathrm{C}_{2} \mathrm{HD}$ band. Hence the band centers may be in error as much as $0.1 \mathrm{~cm} .^{-1}$.
Acknowledgments.-The authors wish to thank Dr. M. Hellman for preparing the sample and Mr. Joseph M. Cameron for performing the least squares reduction on SEAC.
Washingto: 25, D. C.
[Contribution from the Frick Chemical Laboratory, Princeton University, and the Research Laboratories, Eastman Kodak Company Rochester, N. Y.]

# Dipole Moments and Structures of Two Heteropolar Cyanine Dye Molecules ${ }^{1,2}$ 

By Anthony J. Petro, Charles P. Smyth and Leslie G. S. Brooker<br>Received December 29, 1955

The molecular dipole moments of two cyanine dyes have been measured in very dilute solutions in dioxane. Comparison of the observed moment values 9.8 and $13.3 \times 10^{-18}$ with those calculated for the possible resonating structures indicates that a full, or nearly full, electronic charge is carried by the acidic radical of the molecule, while an equal positive charge is equally divided between the nitrogens of the two basic nuclei.

Dipole moment measurements carried out in the Frick Laboratory have been used, ${ }^{3,4}$ in conjunction with absorption data in the visible region, to help clarify the resonance conditions in various nonionized dyes. A long series of papers by Brooker and colleagues dealing with these molecules is listed in reference 4 . Two new dyes are of particular interest because of their possibly large polarities. The two compounds are 1,3 -bis-(2-methoxyethyl)-J-[bis-1,3-(3-ethyl-2-benzothiazolinylidene)-2-pro-pylidene]-barbituric acid and 1,3 -diethyl- 5 -[bis-1,3-(3-ethyl-2 - benzothiazolinylidene)-2-propyli-dene]-2-thiobarbituric acid, both belonging to the group of dyes known as the holopolar cyanines. ${ }^{5}$ They will hereafter be referred to as I and II, respectively.

## Experimental

The dipole moments were measured in dioxane solutions at $30^{\circ}$. "Practical" 1,4-dioxane from the Carbide and

[^1]Carbon Chemicals Corporation was refluxed overnight with sodium and then fractionally distilled as needed. The fraction used boiled at $101.2-101.3^{\circ}$ at 758.0 mm . and had a dielectric constant of $2.2005 \pm 0.0002$ at $30^{\circ}$. Dioxane was chosen as solvent because of the greater solubility of the dyes in this medium than in benzene.

Because of the very slow rate of solution of the dyes, heat was used to dissolve the solids. In the more concentrated solutions, which were evidently supersaturated, the solids recrystallized after two days, so that measurements were carried out as soon as possible after the solutions were prepared.

The dielectric constants of the solutions were measured with a heterodyne-beat apparatus as described elsewhere. ${ }^{3,6}$ The dielectric cell was similar to that previously described ${ }^{7}$ with one modification. Filaments of lead glass (Corning No. 8871) instead of mica strips were used as spacing material between the concentric condenser plates, the glass being fused to the platinum in a high temperature oven. ${ }^{8}$ This construction gave added stability to the cell and thus an invariable cell constant.

Because of the low solubilities of the dyes (range of mole fractions used: I, $(0.5-1.5) \times 10^{-4}$; II, $\left.(1.2-4.4) \times 10^{-5}\right)$, it was considered unprofitable to measure the densities of the more dilute solutions. Consequently, the dersity of only the most concentrated solution was measured in each case, with an Ostwald-Sprengel pycnometer, and a linear dependence of specific volume $\tau_{12}$ on the mole fraction $c_{2}$ was assumed. The values of the slopes $\beta^{\prime}$ in the equation

$$
v_{12}=v_{1}+\beta^{\prime} c_{2}
$$

were then calculated from the specific volumes of the pure solvent and the most concentrated solution. The effect of

[^2]
[^0]:    (1) The work reported herein was supported by the U. S. Atomic Energy Commission.
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[^1]:    (1) This research was supported by the United States Air Force through the Office of Scientific Research of the Air Research and Development Command. Reproduction, translation, publication, use or disposal in whole or in part by or for the United States Govern, ment is permitted.
    (2) This paper represents a part of the work to be submitted by Mr. A. J. Petro to the Graduate School of Princeton University in partial fulfilment of the requirements for the degree of Doctor of Fhilosophy.
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